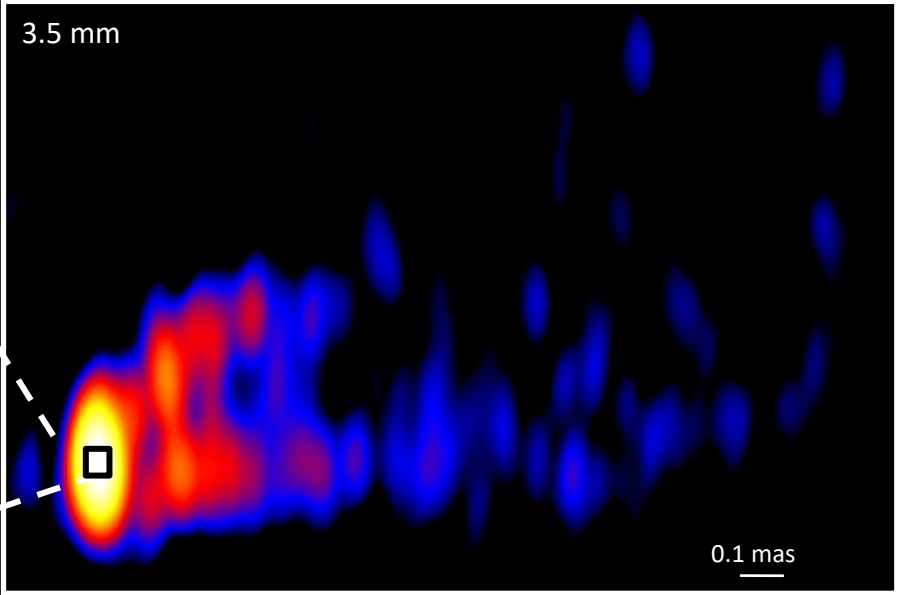
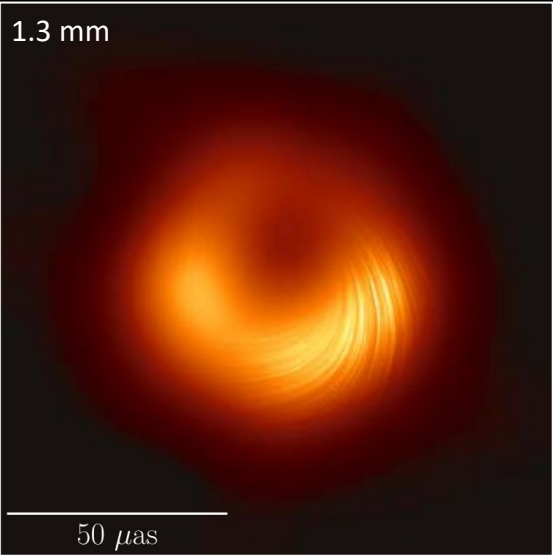


# Polarization Signatures of Jet Launching on Horizon Scales

Andrew Chael  
Princeton Gravity Initiative  
5/24/24



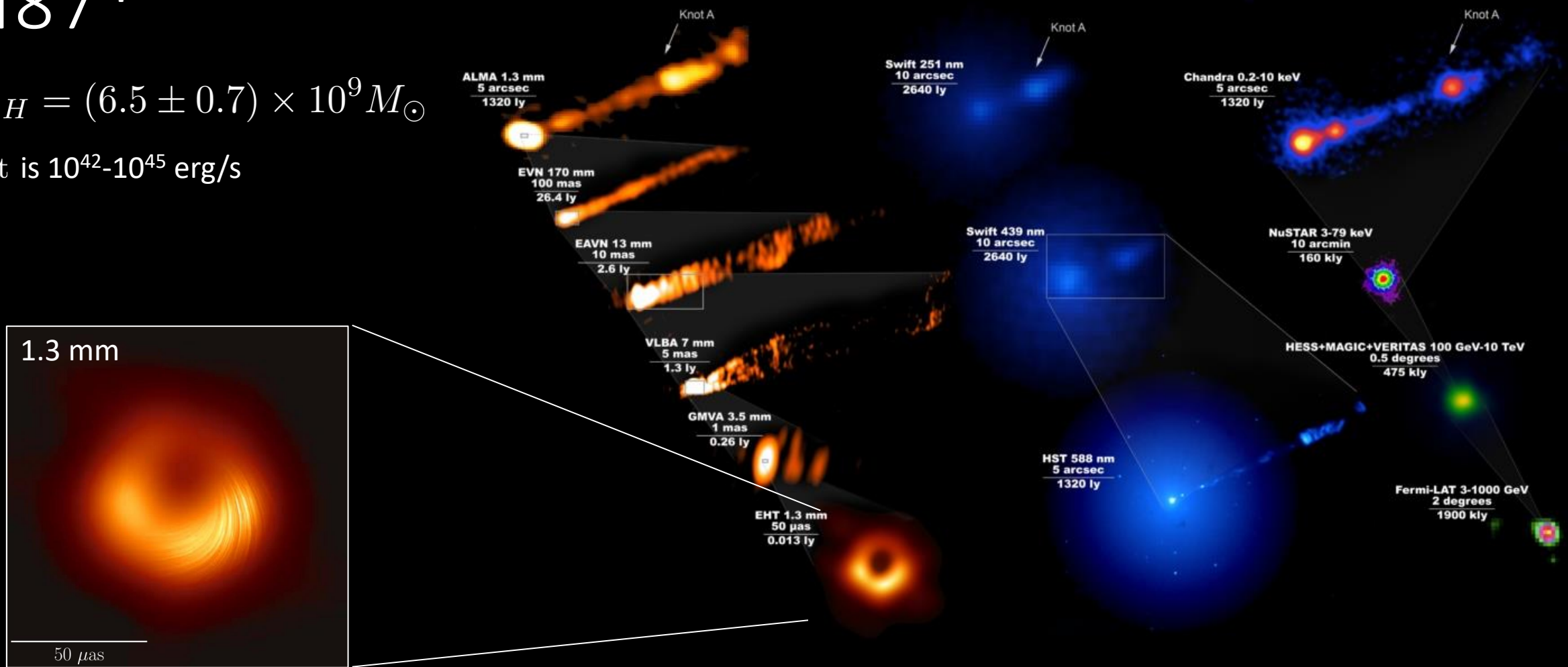
Event Horizon Telescope

Image Credits: Kazuhiro Hada (VLBA+GBT 3mm), EHT (1.3 mm)

# M87\*

$$M_{BH} = (6.5 \pm 0.7) \times 10^9 M_{\odot}$$

$$P_{jet} \text{ is } 10^{42}\text{-}10^{45} \text{ erg/s}$$



Jets are thought to be powered by black hole spin energy extracted via magnetic fields (Blandford & Znajek 1977)  
Is it possible to observe black hole energy extraction **on horizon scales**?

# M87's Jet in Simulations

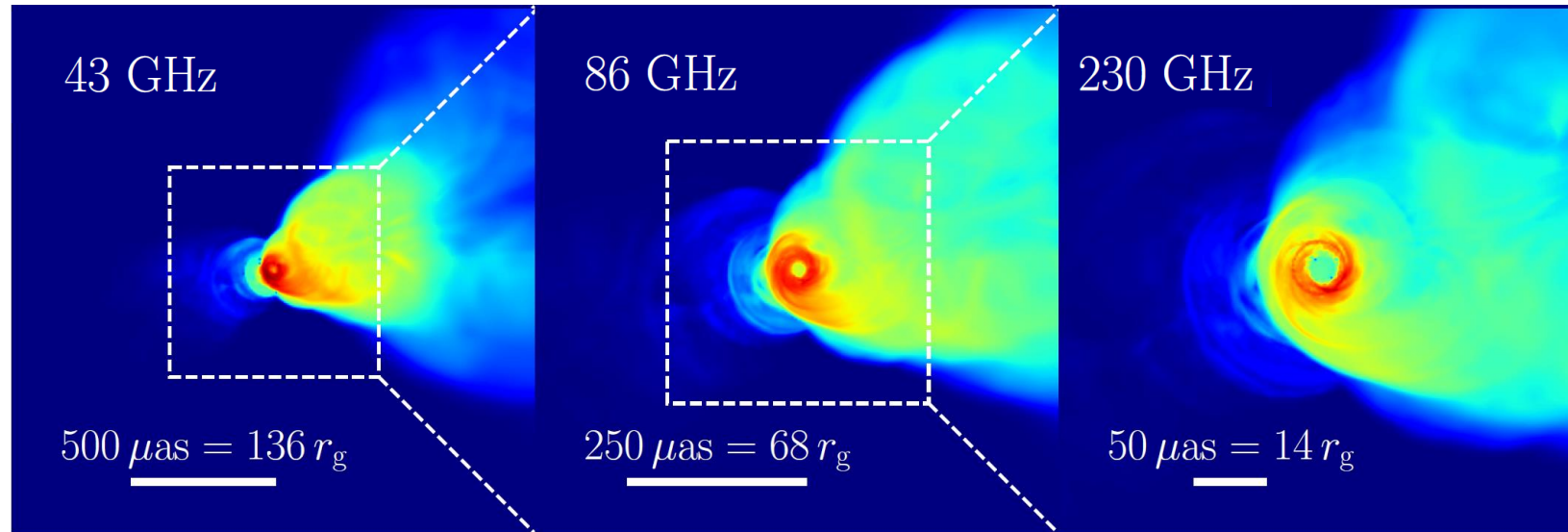
- Jets from General Relativistic Magnetohydrodynamic (**GRMHD**) simulations **are powered by black hole spin**

(e.g. McKinney & Gammie 2004, Tchekhovskoy+ 2012, EHTC+ 2019, Narayan+ 2022)

- Radiative GRMHD simulations naturally produce the correct:

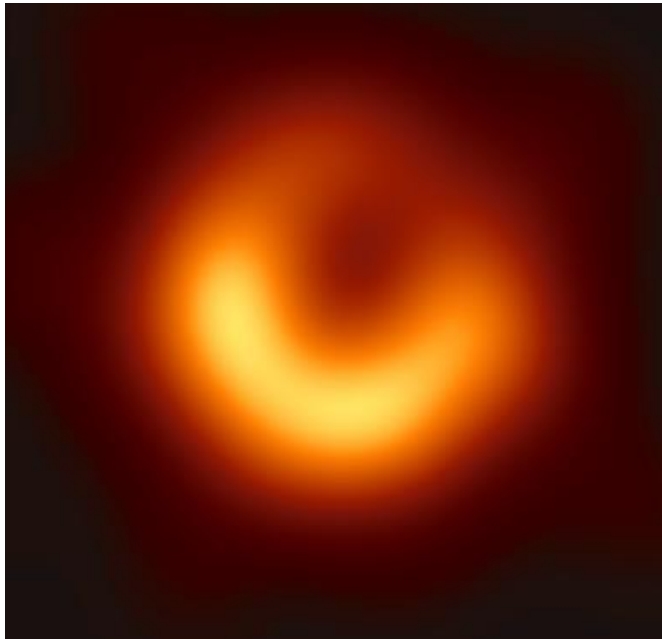
- jet power
- wide opening angle
- core-shift

- Can we be **sure** the jet is BZ?  
What is a **physically meaningful** observation of **horizon-scale** energy flow?

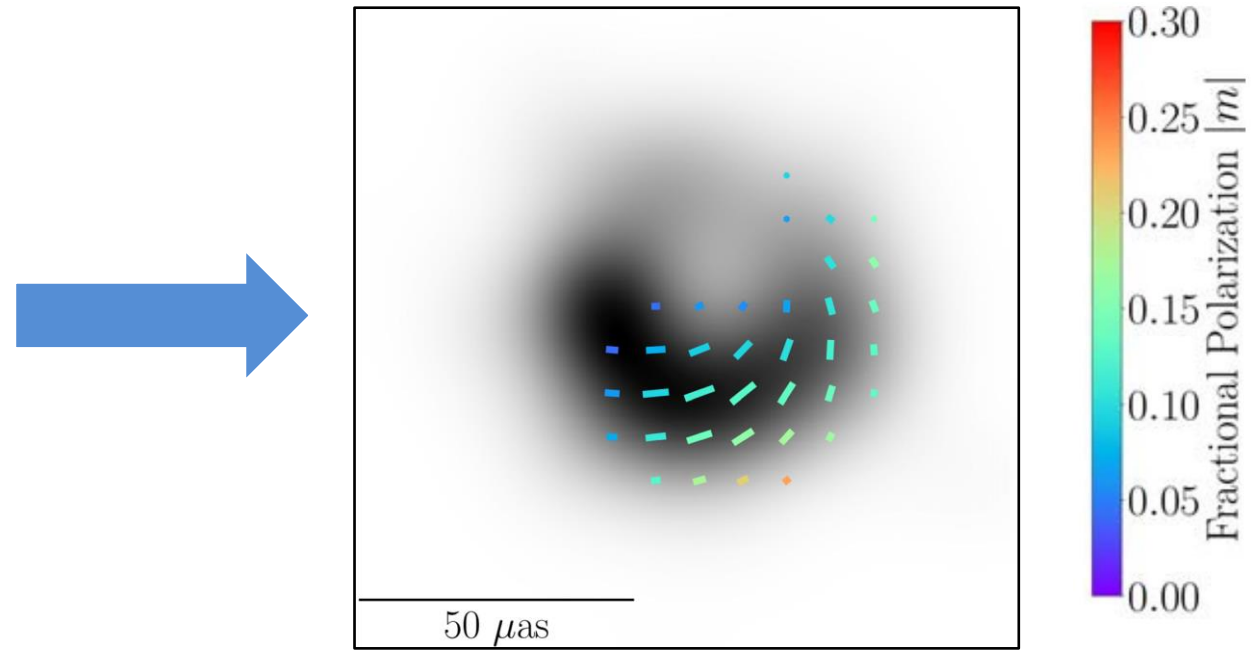


# M87\* 2017 in linear polarization

Total intensity

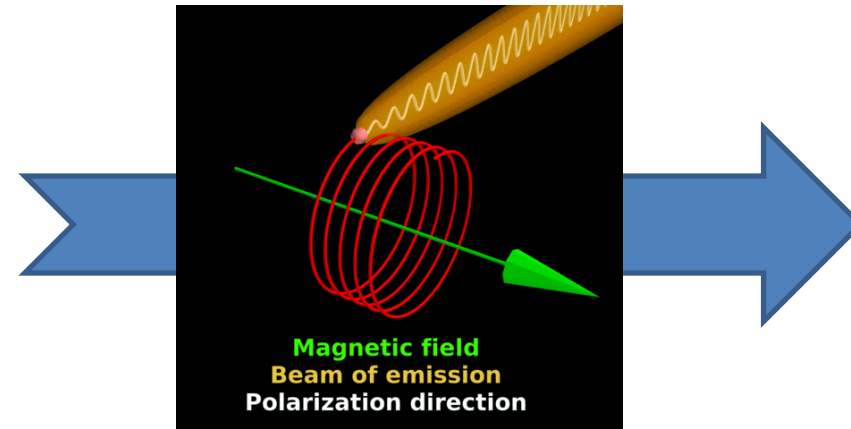
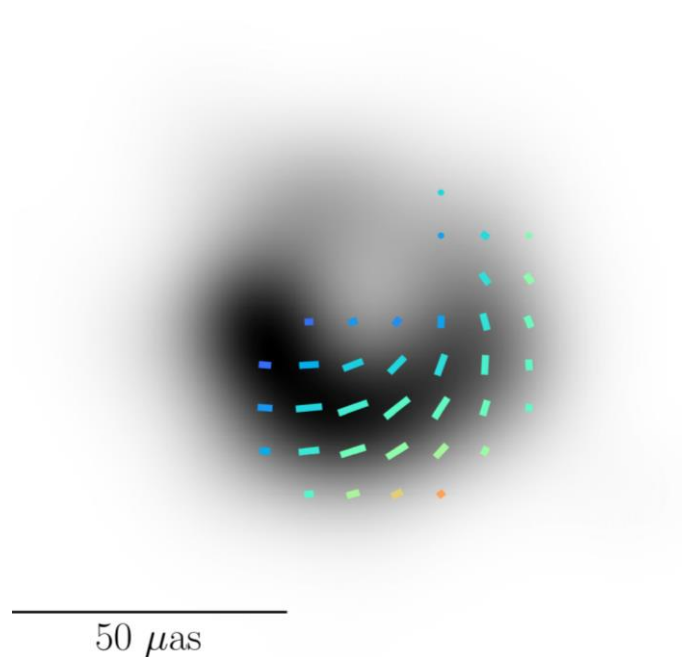


Linear Polarization



- Polarization is concentrated in the southwest
- Polarization angle structure is predominantly **helical**
- Overall level of polarization is **weak**,  $\sim 15\%$

# Why polarization?

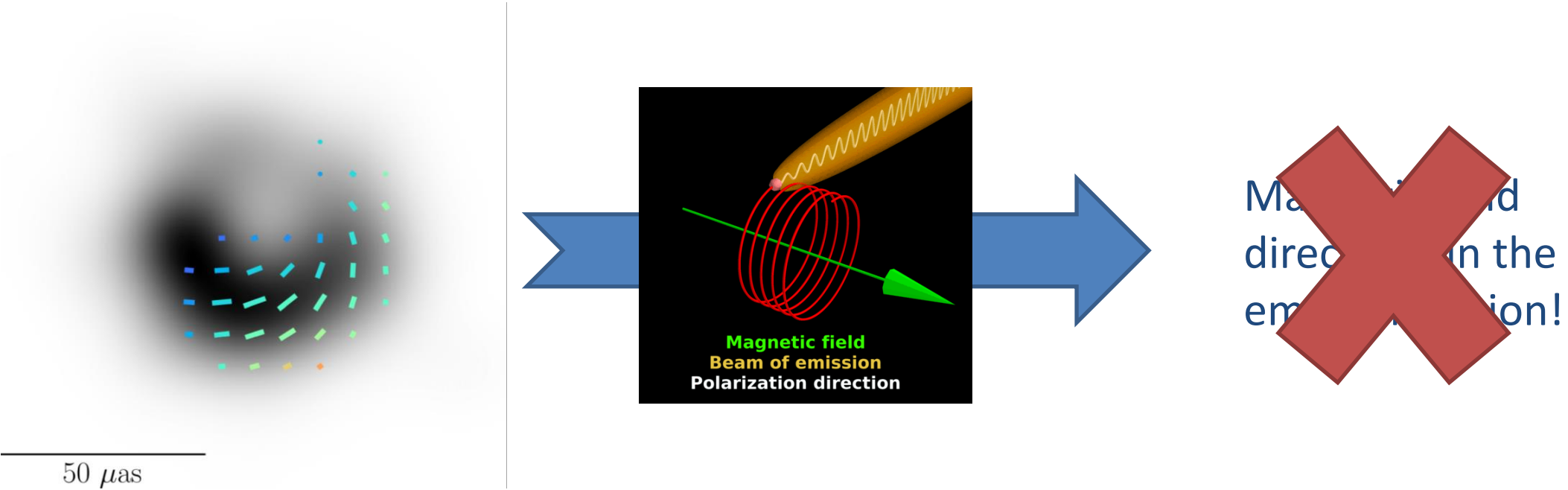


Magnetic fields in  
the emission  
region!

Synchrotron radiation is emitted with polarization **perpendicular** to magnetic field lines

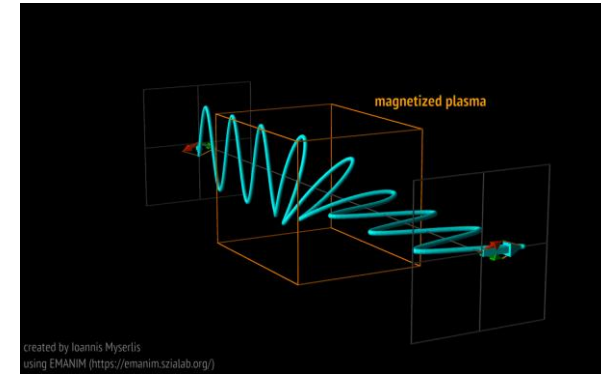
Polarization **transport** is sensitive to the magnetic field, plasma, and spacetime

# Synchrotron polarization traces magnetic fields?

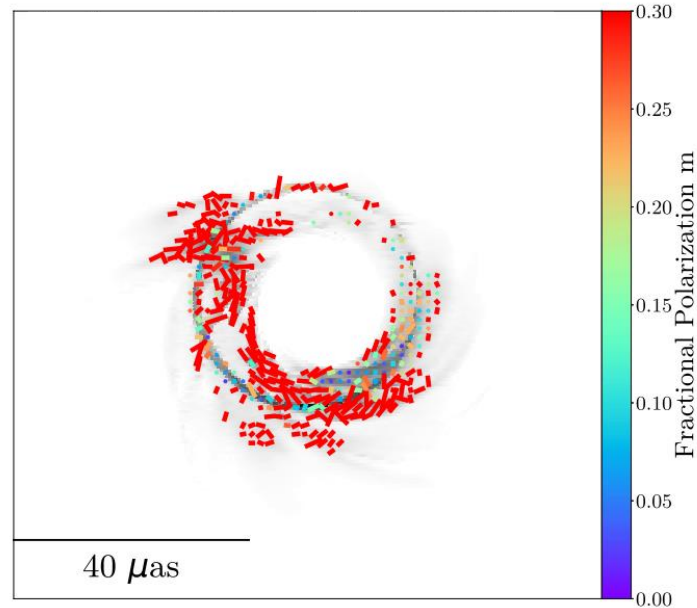


**GR and Faraday effects make the situation in M87\* more complicated!**

# Faraday Rotation is important!

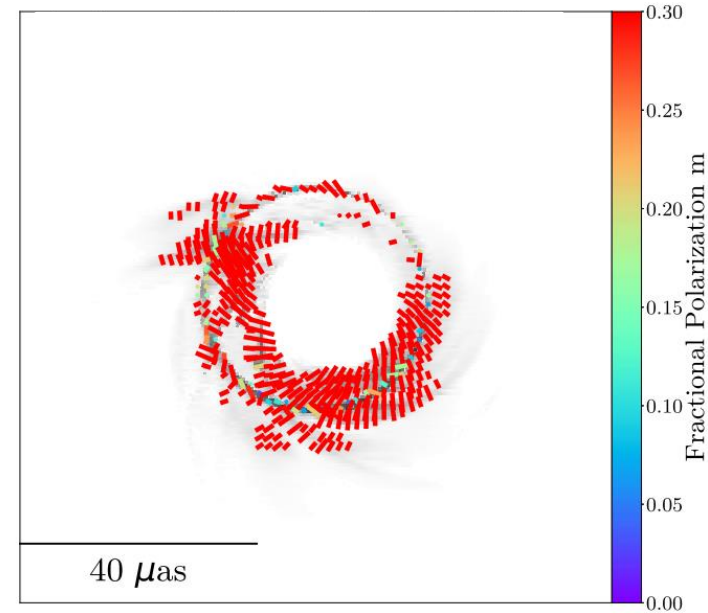


With rotation



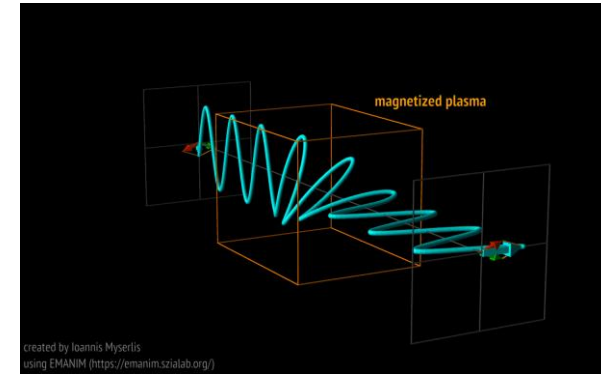
'infinite' resolution

Without rotation

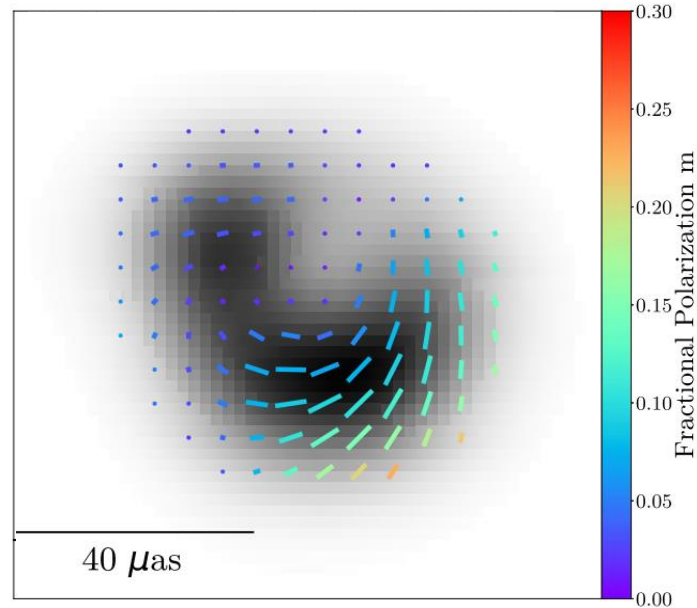


- Significant Faraday rotation on small scales  
→ **scrambles** polarization directions

# Faraday Rotation is important!

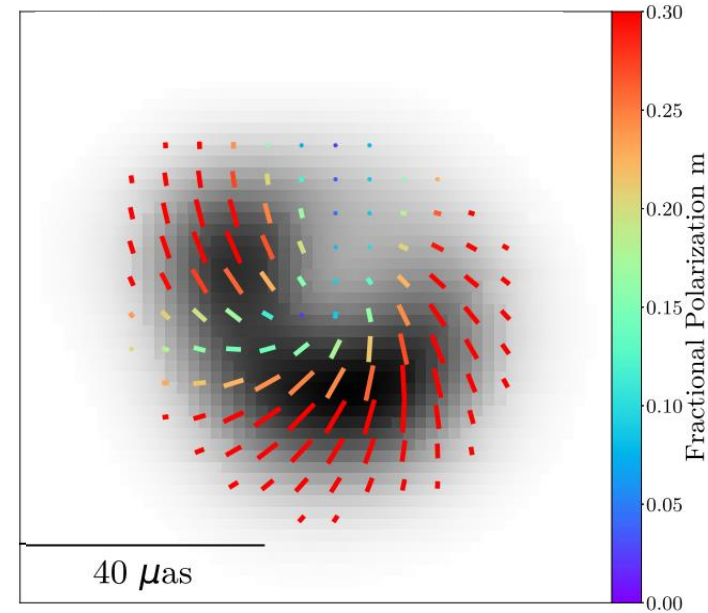


With rotation



EHT resolution

Without rotation



- Significant Faraday rotation on small scales
  - **scrambles** polarization directions
  - **depolarization** of the image when blurred to EHT resolution
  - **rotates** the pattern when blurred to EHT resolution



# Scoring simulations with polarization: Results

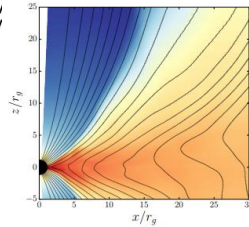
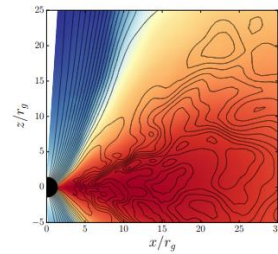
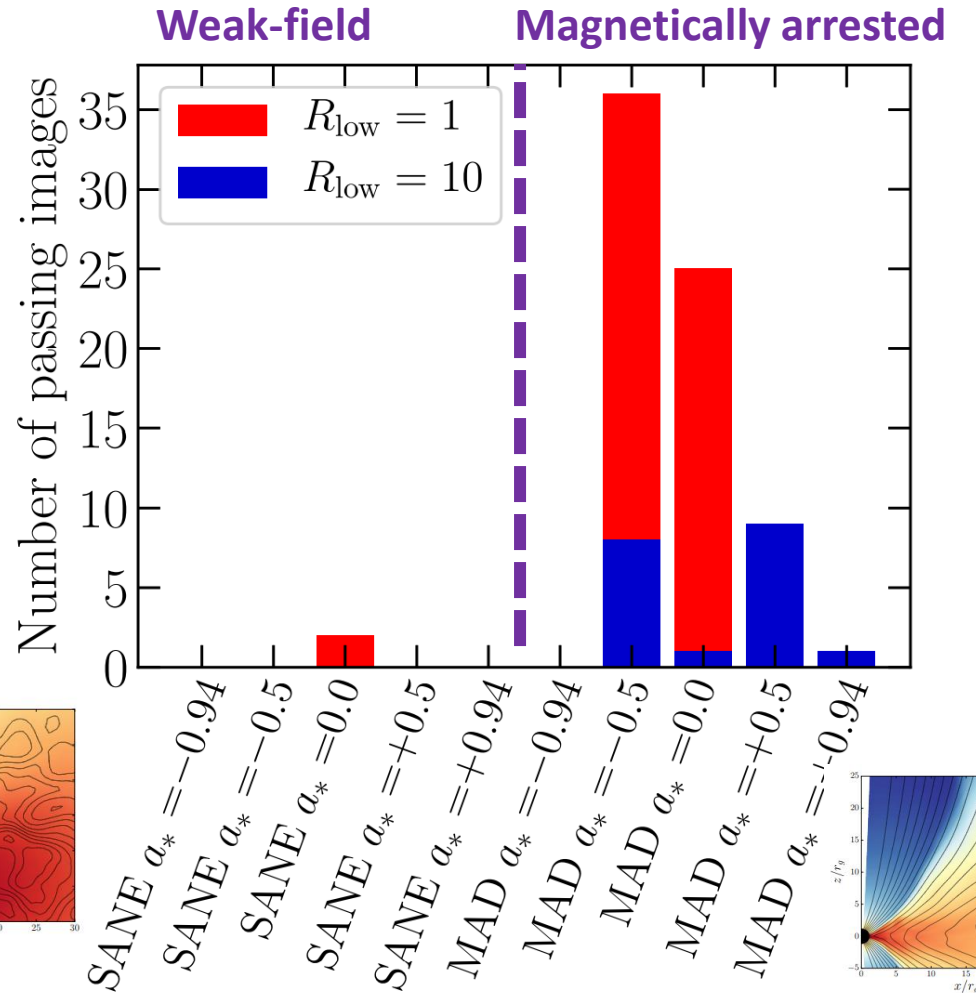
- Scoring with multiple approaches **all strongly favor a magnetically arrested accretion flow**

- We constrain M87\*'s allowed accretion rate by 2 orders of magnitude:

$$\dot{M} \simeq (3 - 20) \times 10^{-4} M_{\odot} \text{ yr}^{-1}$$

$$\left( \dot{M}_{\text{Edd}} = 137 M_{\odot} \text{ yr}^{-1} \right)$$

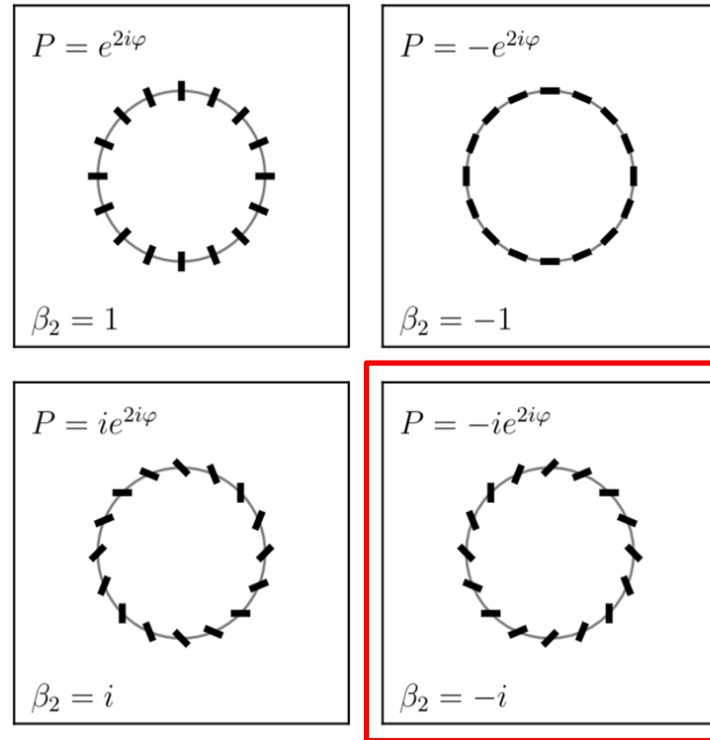
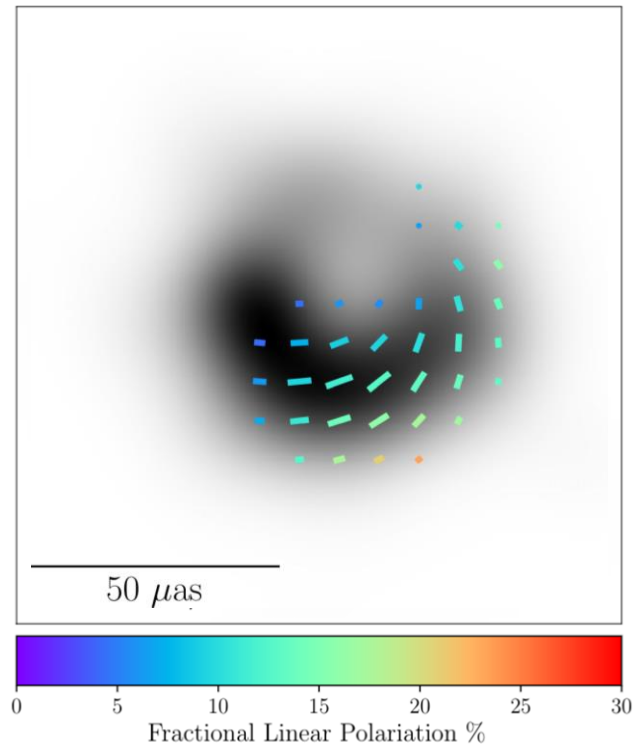
- Strong fields **more easily launch jets** at lower values of BH spin



Can we connect the polarized image of M87\*  
on horizon scales to energy flow & jet  
launching?

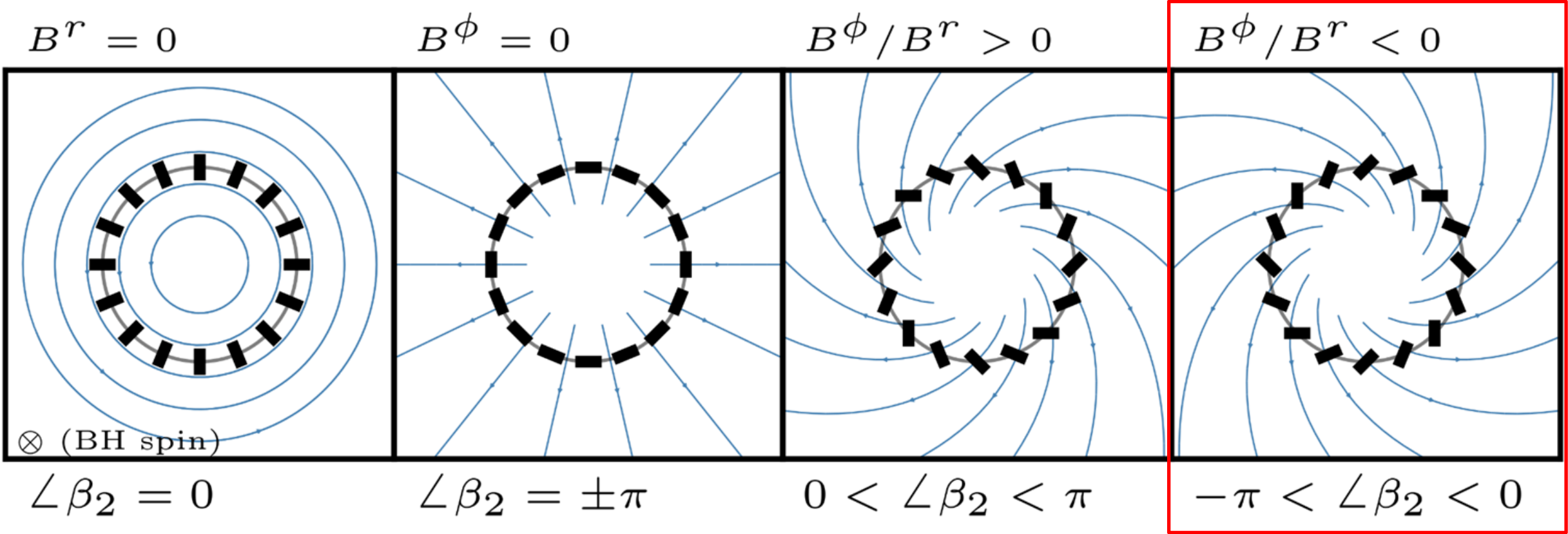
Chael, Lupsasca, Wong & Quataert 2023  
[arXiv: 2307.06372](https://arxiv.org/abs/2307.06372)

# Polarized Images of M87\* and horizon-scale energy flow



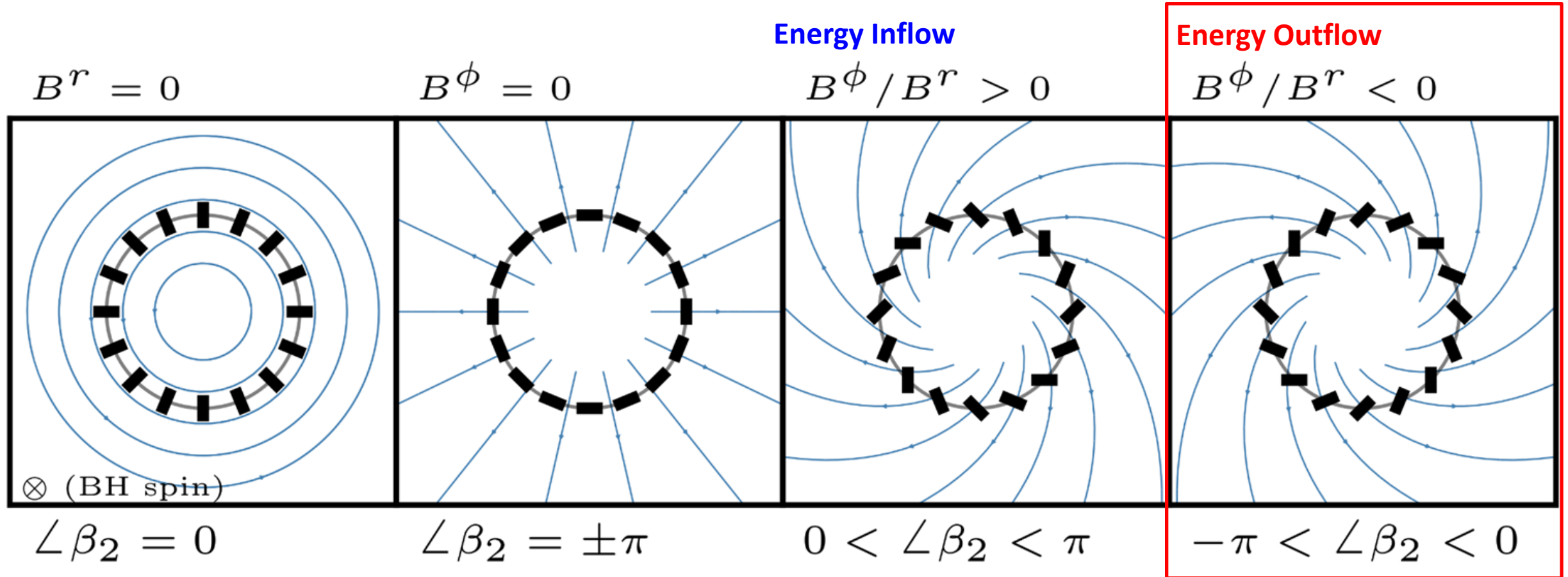
- The polarization spiral's **2<sup>nd</sup> Fourier mode** ( $\beta_2$ : Palumbo+ 2020) is the **most constraining** image feature
- Can we interpret  $\beta_2$  **physically**?

Cartoon model:  $\arg(\beta_2)$  is connected to the pitch angle  $B^\phi / B^r$



- Face on fields, no Faraday rotation, no optical depth, no relativity
- Coordinate axis is **into the screen/sky** (EHT Paper V, 2019)

$\arg(\beta_2)$  is connected to the **electromagnetic energy flux**



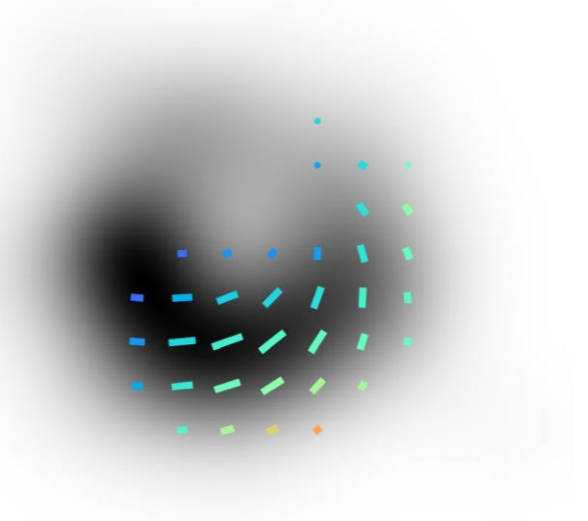
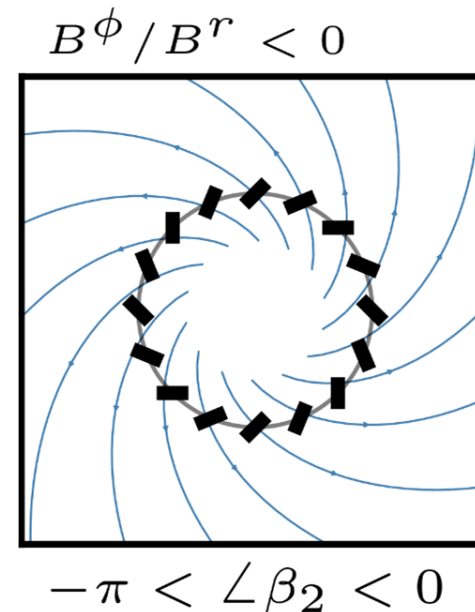
**Radial Poynting flux in Boyer-Lindquist coordinates:**

$$\mathcal{J}_E^r = -T_{t \text{ EM}}^r = -B^r B^\phi \Omega_F \underbrace{\Delta \sin^2 \theta}_{\geq 0}$$

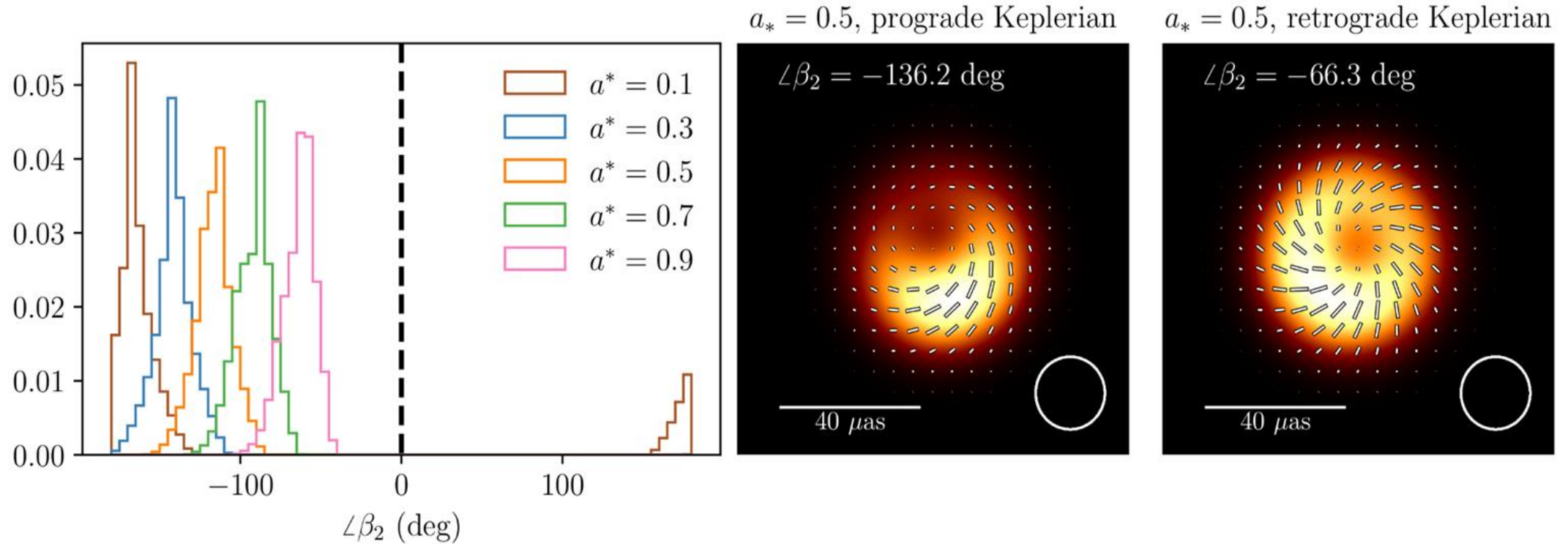
↑  
fieldline angular speed

# $\arg(\beta_2)$ is connected to the electromagnetic energy flux

- The sign of  $\arg(\beta_2)$  is directly connected to the direction of Poynting flux, assuming we know the sign of  $\Omega_F$
- Ignoring Faraday effects, **the EHT's measurement of  $\beta_2$  implies electromagnetic energy is outflowing in M87\***
- This inference requires we assume fieldlines **co-rotate** with the emitting plasma in a **clockwise** sense
- Does this simple argument hold up in **more complicated models**?



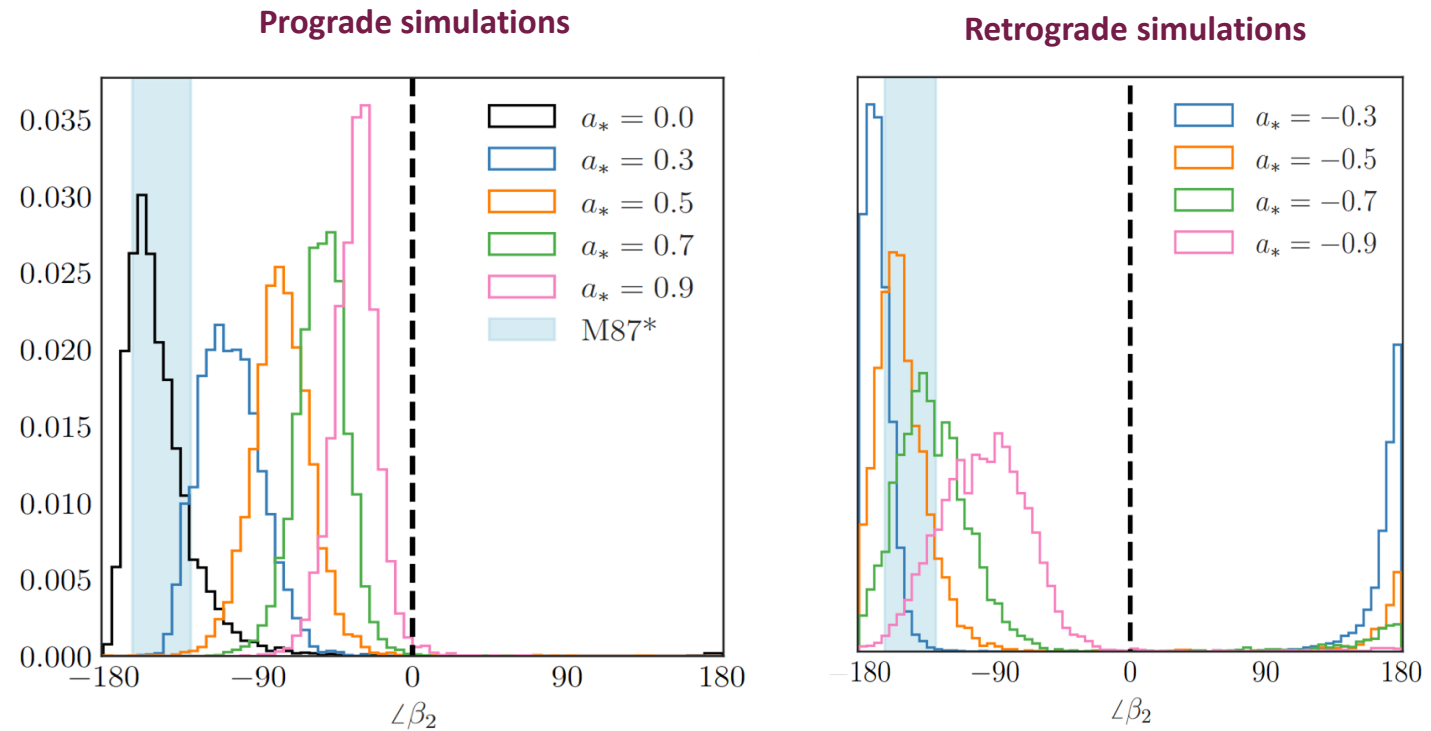
# $\arg(\beta_2)$ in semi-analytic models of M87\*



- We fix magnetic fields to the BZ monopole solution (with energy outflow)
- The black hole spin direction is fixed into the sky
- We explore many models for the velocity of the emitting fluid

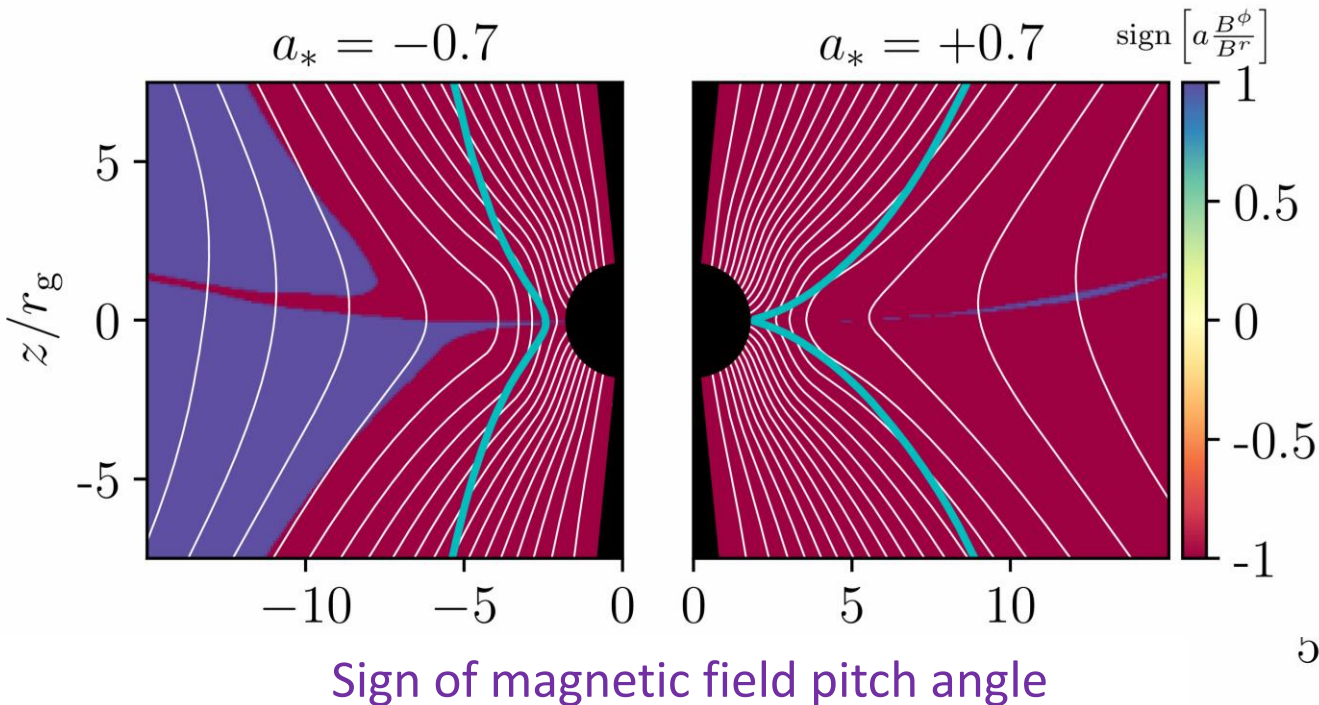
# $\arg(\beta_2)$ in MAD GRMHD simulations of M87\*

- 1600 simulated EHT-resolution M87\* images from MAD simulations (Narayan+ 2022)
- Almost all 230 GHz simulation images have **negative**  $\arg(\beta_2)$  consistent with the measured energy outflow in the simulations
- $\arg(\beta_2)$  has the **same qualitative dependence on spin** as in a simple BZ monopole model!

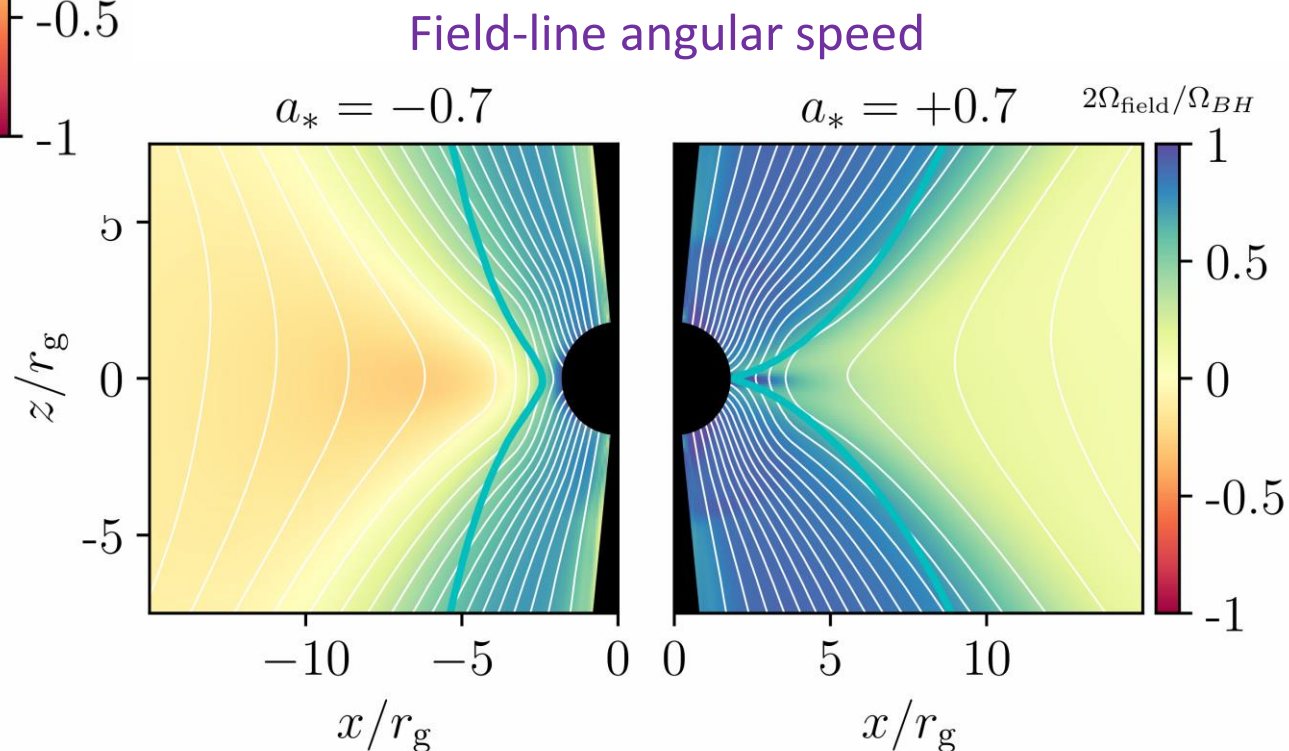




In GRMHD, energy-extracting fieldlines set  $\arg(\beta_2)$

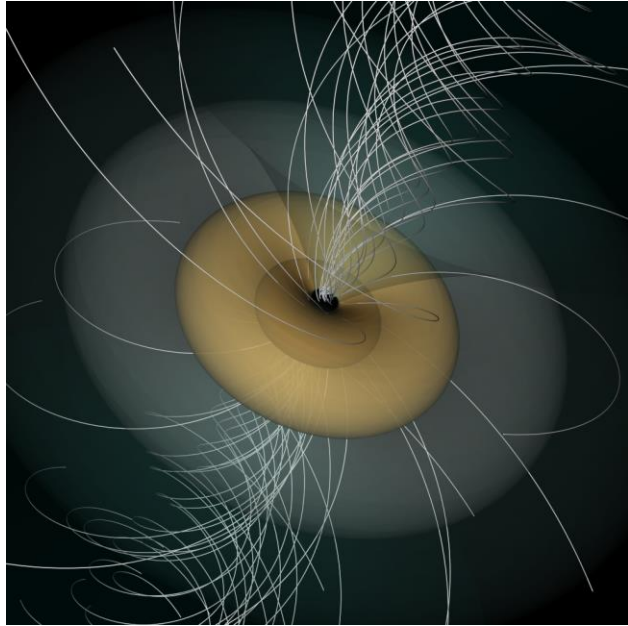


Even in **retrograde** simulations, field-lines in the 230 GHz emission region usually corotate with the black hole and have a negative  $B^\phi / B^r$

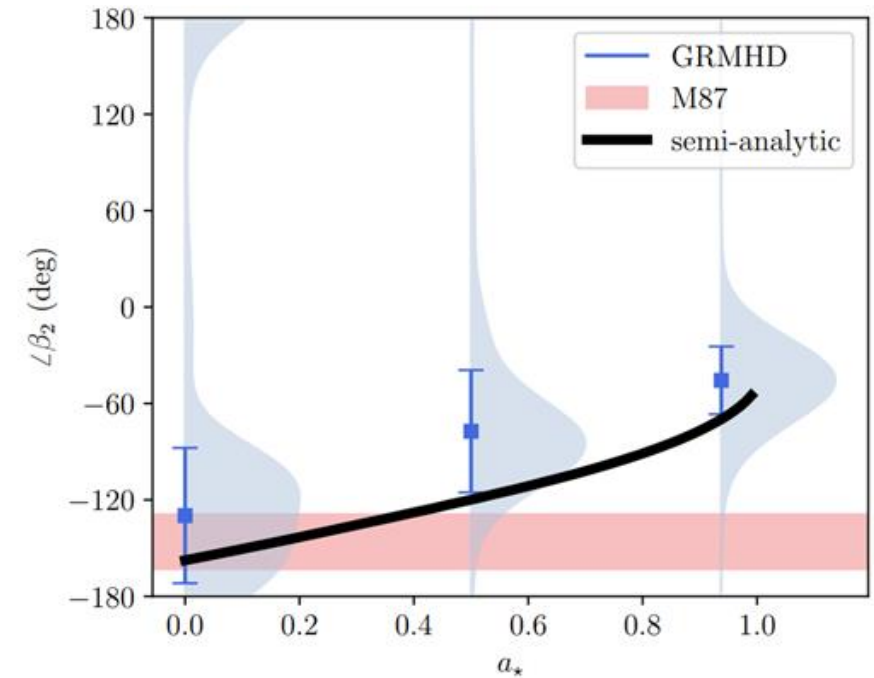
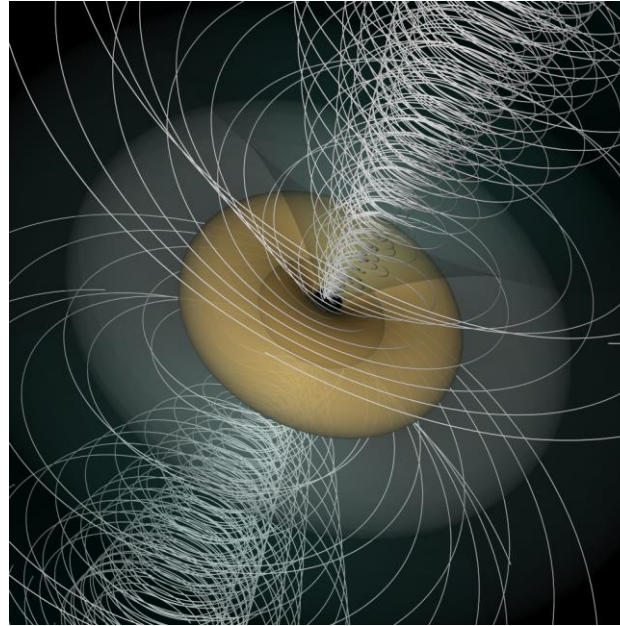


# Polarized images are **spin dependent**

Low Spin

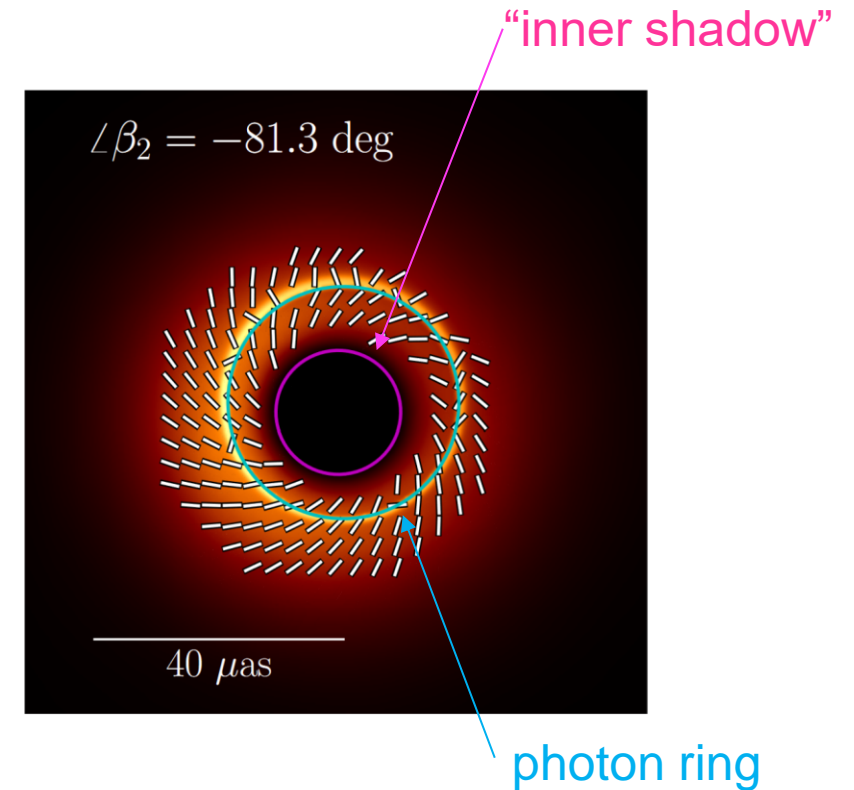
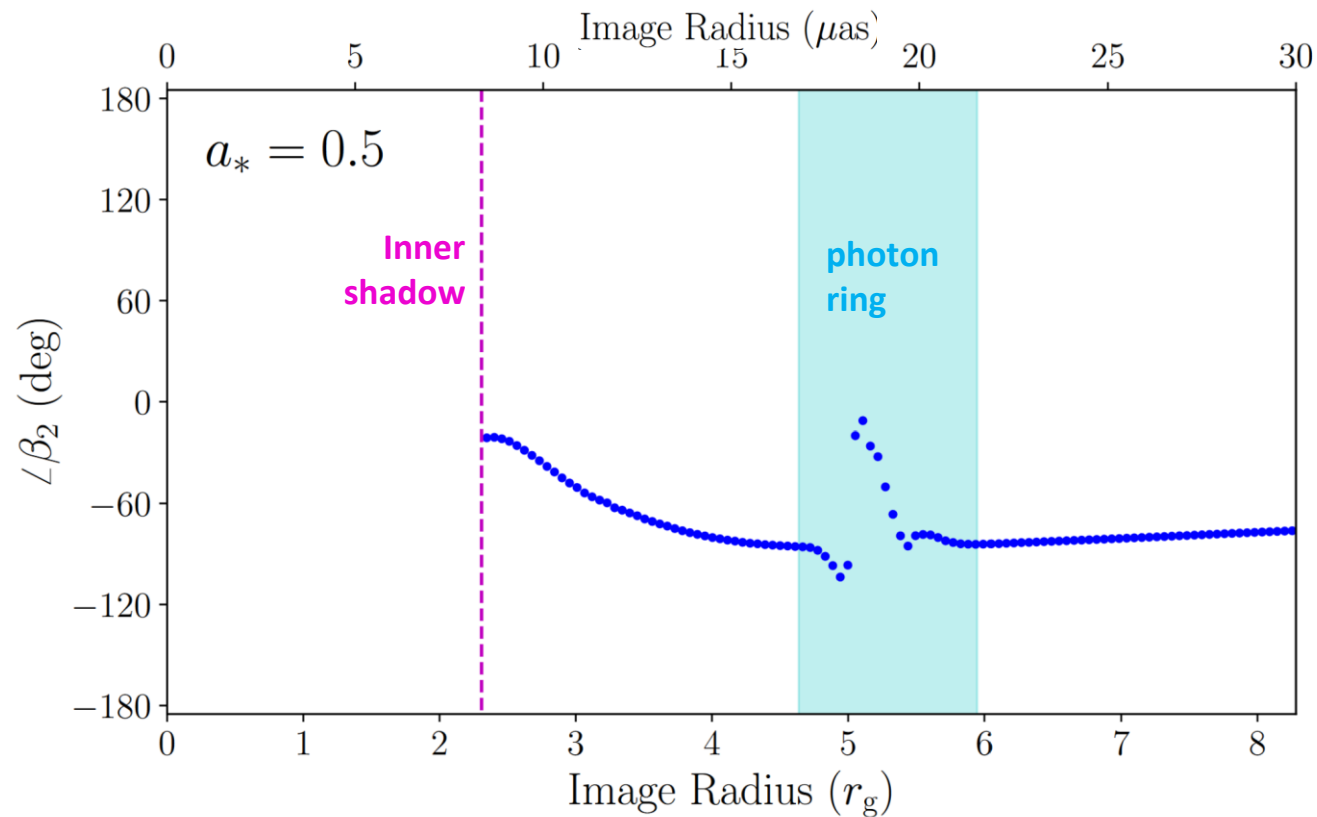


High Spin



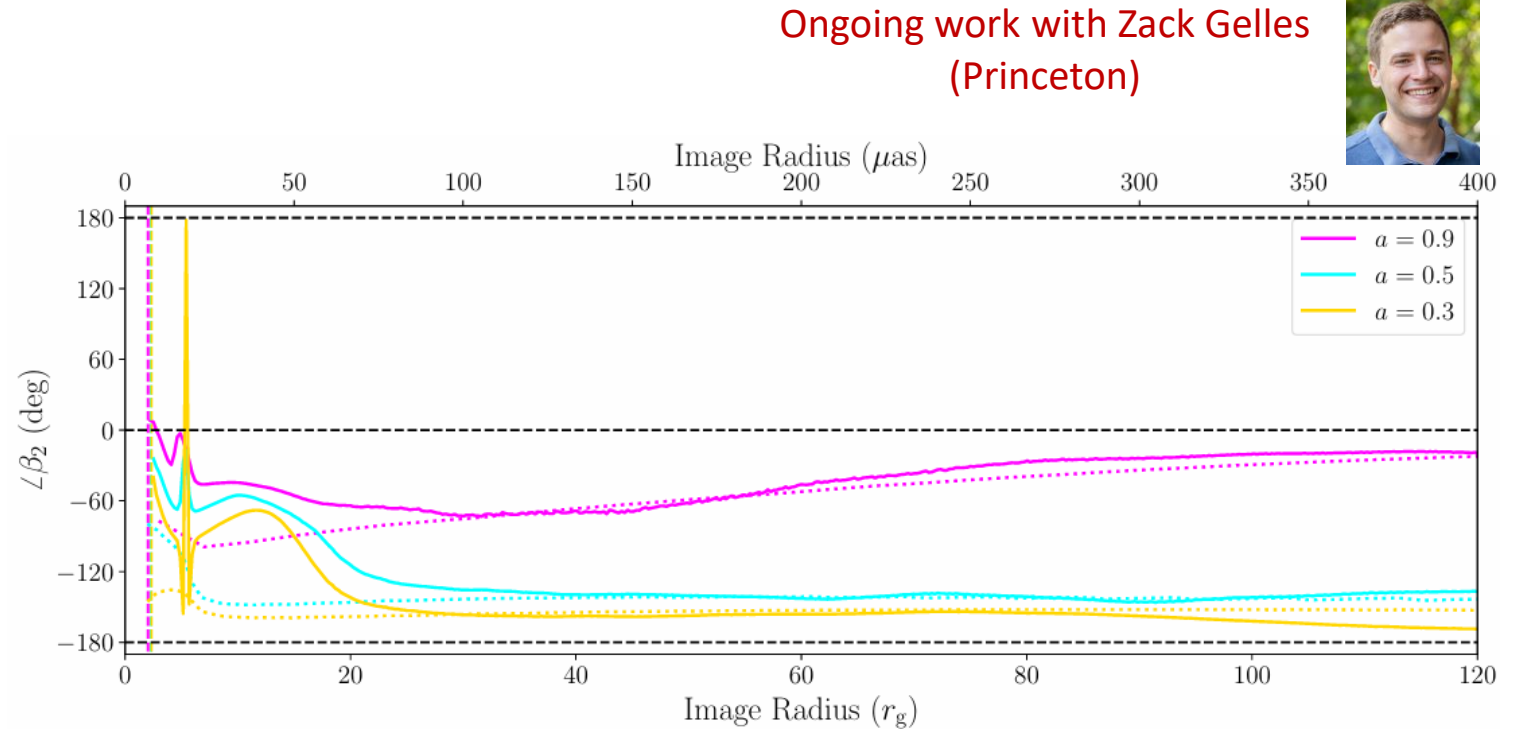
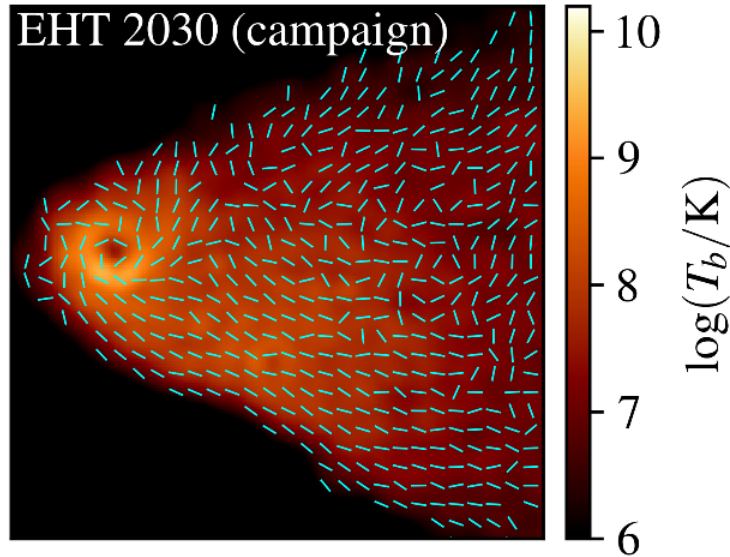
- Black hole **spin winds up initially radial fields**, but always so that  $B^\phi / B^r < 0$
- The field pitch angle **increases with spin**
- Increased field winding
  - increases the BZ jet power
  - and makes the observed polarization pattern more radial

# To look for energy extraction, we need to zoom in



- Measuring polarization as a function of radius **probes energy flow at different scales**
- Both simple models and GRMHD simulations make a strong prediction
  - $\arg(\beta_2)$  evolves rapidly close to the horizon as the rest frame fields become more azimuthal from **GR frame dragging**

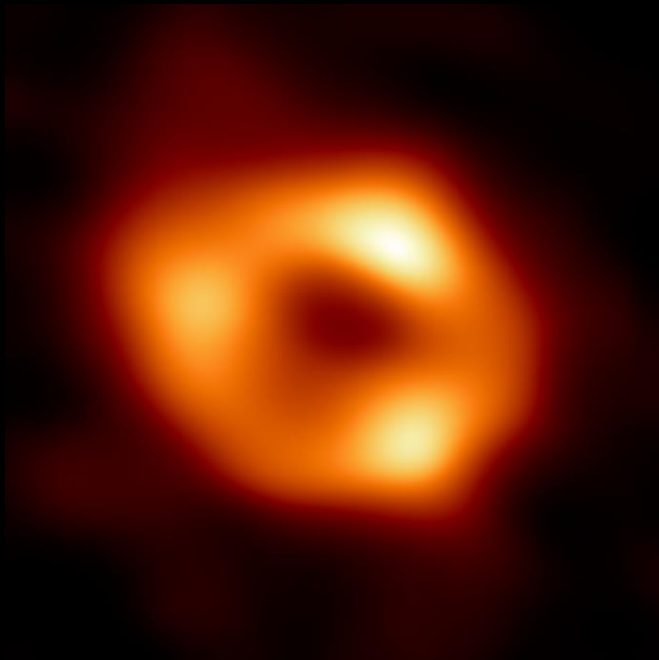
# To look for energy extraction, we need to zoom out



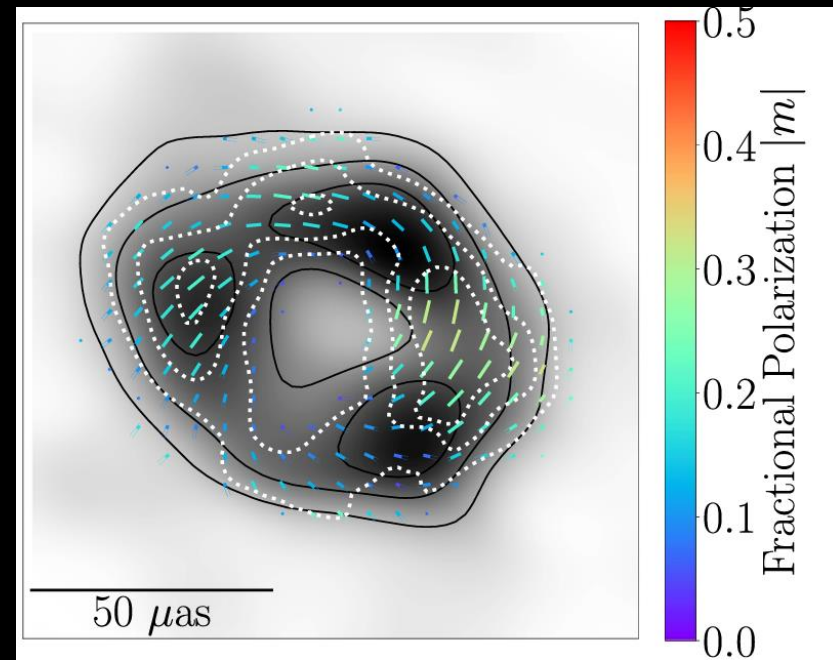
- New telescope sites & larger bandwidth will enhance EHT's **dynamic range**
  - These will illuminate both the **BH-jet connection**
- These new observations will require new theoretical models and simulations to fully interpret
  - Can we directly measure energy flow **from the horizon through the jet base?**

# Sgr A\* in linear polarization

Total intensity



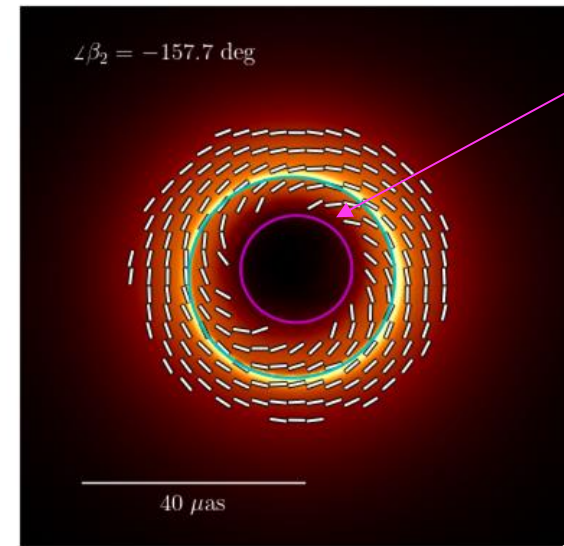
Linear Polarization



- Polarization fraction is **higher** than M87
- $\beta_2$  is consistent with **clockwise rotation** measured in NIR flares
  - **only after Faraday derotation**
- MAD simulations preferred – **where is the jet?**

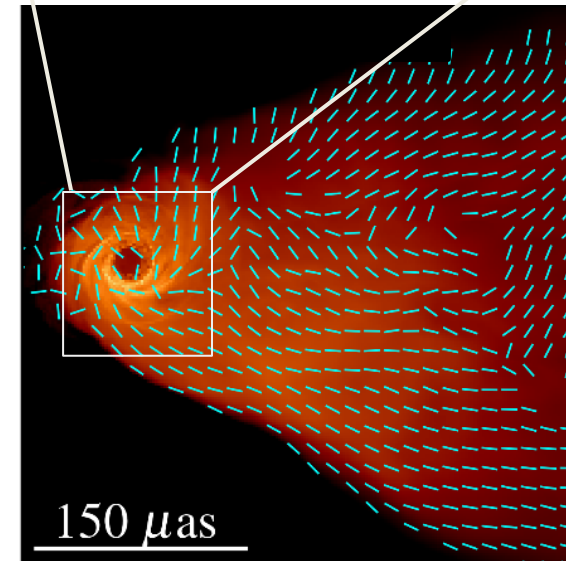
# Takeaways

- Testing the BZ mechanism and directly measuring BH spin in M87\* and other sources is a key science goal for the EHT's next decade
- We need **high-dynamic range, multi-frequency, polarized EHT images** to:
  - Measure polarization **down to the horizon**
  - Connect the energy flux **from horizon scales out through the jet base**
- We need **new simulation suites and analytic models** to
  - calibrate the spin-dependence of  $\beta_2$
  - fully account for complicating factors (Faraday effects, field geometries, 3D structure...)
- Connecting theory and observation, we will be able to directly test the BZ mechanism for jet launching in the next several years



"inner shadow"

Goal 1:  
measure  
energy flux  
**down to  
horizon**



Goal 2:  
measure  
energy flux  
**out through  
jet base**